BLACK HOLE WEATHER

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MS0735.6 cluster (McNamara+05)

SCALING RELATION: HOT HALO - SMBHS



X-RAY TEMPERATURE

HOT HALO SCALING RELATIONS OF SMBHs



X-ray scalings lower scatter + larger corr than optical scalings!

Gaspari et al. 2019



X-RAY HALO SCALING RELATIONS OF SMBHs

(indirect properties, within core)

• gas mass/Yx tighter than total mass





GOAL 1: first-principle multi-scale simulations GOAL 2: test detailed synthetic models with multi- λ data

MG+2011-2020

"BLACK HOLE WEATHER" PROGRAM



multiphase condensation cascade ("raining")

CCA = CHAOTIC COLD ACCRETION

BLACK HOLE FEEDING WITH COOLING: 3 DYNAMICAL STAGES



RAINING ON BLACK HOLES

a.k.a. Chaotic Cold Accretion [CCA] — Gaspari et al. 2013



chaotic streamlines => recurrent
 multiphase gas interactions



TURBULENCE > ROTATION

 $\sigma_v \sim 150 \text{ km/s}$ as found by *Hitomi*

RGB surface density: plasma (blue), warm gas (red), cold gas (green)

CCA has been corroborated by several independent observational and theoretical/simulation studies: e.g., Voit & Donahue 2015, Voit 2015, 2017, 2018; Werner+2014; David+2014, Li & Bryan 2014, 2015; Wong+2014; Russell+2015; Valentini & Brighenti 2015; Yang+2015-2016; Meece+2016; Tremblay+2015, 2016, 2018; Prasad+2016; David+2017; McDonald+2018; Maccagni+2018; Nagai+2019; Rose+2019-2020; Storchi-Bergmann+2019 (review); Schellenberger+2020, ...

TOP-DOWN MULTIPHASE GAS CONDENSATION RAIN: synthetic imaging



Gaspari et al. 2017

condensation criterion:

KINEMATIC TRACERS MULTIPHASE RAIN

Gaspari et al. 2018

ENSEMBLE beam (R < 50 kpc ~ arcmin)

spectral line broadening
= turbulent motions



self-regulated AGN jet feedback sims

global turbulence kinematics: ensemble warm phase and hot/plasma phase are linearly related

similar can be shown for UV - IR - radio (molecular) phases:

multiwavelength synergies: ATHENA - ALMA - JWST/ELT - VLT/MUSE, SINFONI - SKA

CHAOTIC COLD ACCRETION DYNAMICS

Gaspari+2017



- leaf clouds via clump finder algorithm
- network of condensed structures
- key for AGN obscuration/unification models (BLR, NLR)
- angular momentum mixing/cancellation via inelastic collisions

can be modeled as quasi-spherical viscous accretion:

$$\begin{split} \lambda_{\rm c} &\equiv \frac{1}{n_{\rm c} \, \pi (2 \, r_{\rm c})^2} = \frac{1}{3} \frac{r_{\rm c}}{f_V} \simeq 88^{+262}_{-67} \, \rm pc \qquad \text{mean free path} \\ \nu_{\rm c} &\equiv \sigma_v \, \lambda_c \simeq 4.5^{+13.3}_{-3.1} \times 10^{27} \, \rm cm^2 \, s^{-1} \qquad \text{effective collisional viscosity} \\ \dot{M}_{\bullet} &= 4.8 \times 10^{-3} \, \nu_{\rm c} \simeq 0.3^{+0.9}_{-0.2} \, \rm M_{\odot} \, yr^{-1} \qquad \text{average inflow rate} \\ \text{(for massive ETG)} \end{split}$$

recurrent 2 dex boost in accretion rate ~ 100x Bondi rate



CHAOTIC COLD ACCRETION VARIABILITY



constant variance per log interval => large self-similar variability on different timescales

characteristic of fractal and chaotic phenomena:

quasars (e.g., 3C273), sunspots, meteorological data/RAINFALLS, heart beat rhythms, neural activity, stock market, ...

BLACK HOLE FEEDING & FEEDBACK SELF-REGULATION





Gaspari+12



X-RAY BUBBLES



NGC 5813 - Randall+2011



COCOON SHOCKS

7.1

7.0

6.9

MS0735.6 cluster (McNamara+2005)



MG+2012b SIMS

$$\Pi_{\rm s} = \frac{(\gamma+1)}{12\gamma^2} \frac{\omega p}{2\pi} \left(\frac{\delta p}{p}\right)^3 \label{eq:sigma_s}$$

weak shock heating

TURBULENCE

"Enstrophy" ~ magnitude (squared) of vorticity/turbulence



- stretching motions balanced by rarefactions
- baroclinic motions subdominant

Lagrangian tracers on top of AMR hydro simulation





Gaspari+20 *Nature Ast.* review