

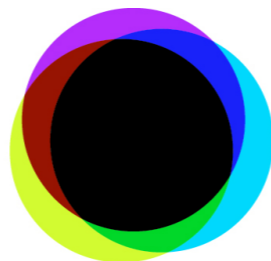
A new era in the quest for Dark Matter

Gianfranco Bertone

GRAPPA center of excellence, U. of Amsterdam

Seminar @ Astrophysics and Space Science Observatory, Bologna 22/6/2021

GRAPPA x
x
x



GRavitation AstroParticle Physics Amsterdam



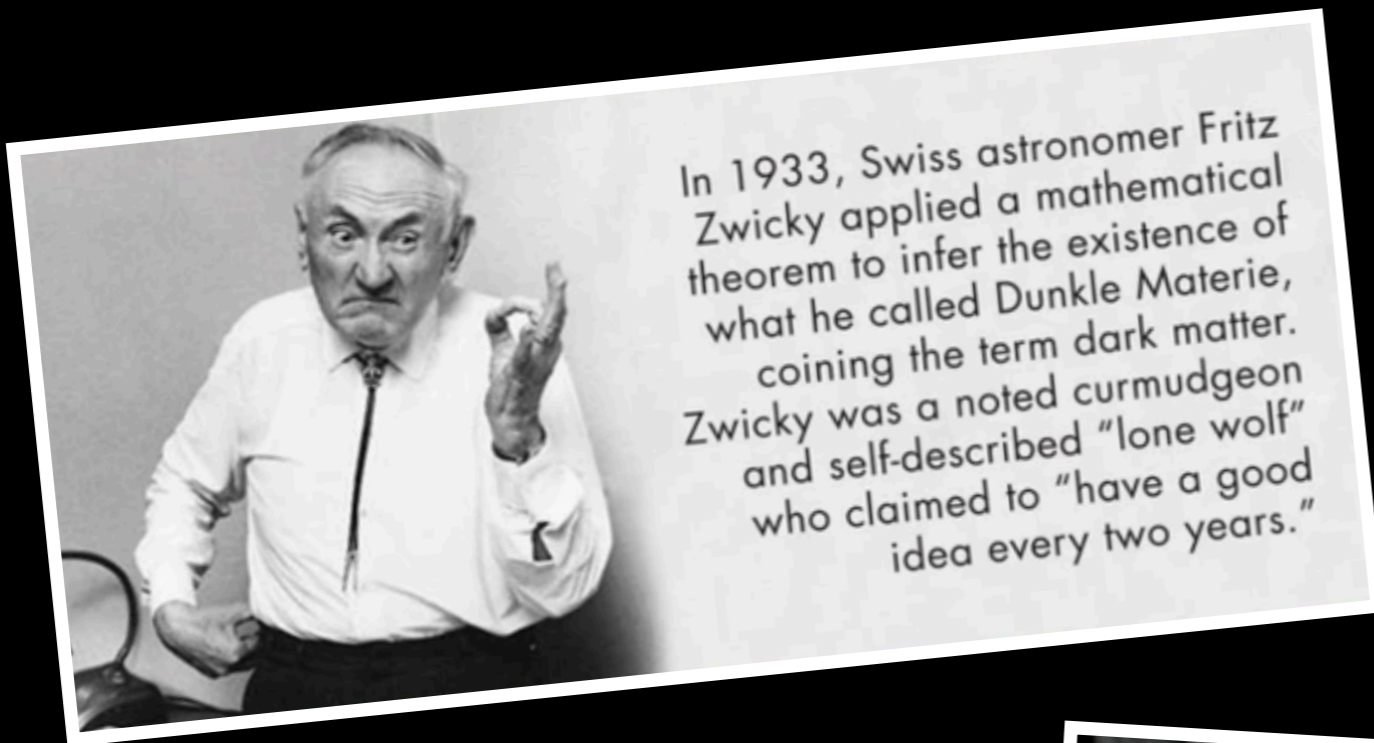
Plan of the talk:

Preamble: the dark universe *narrative*

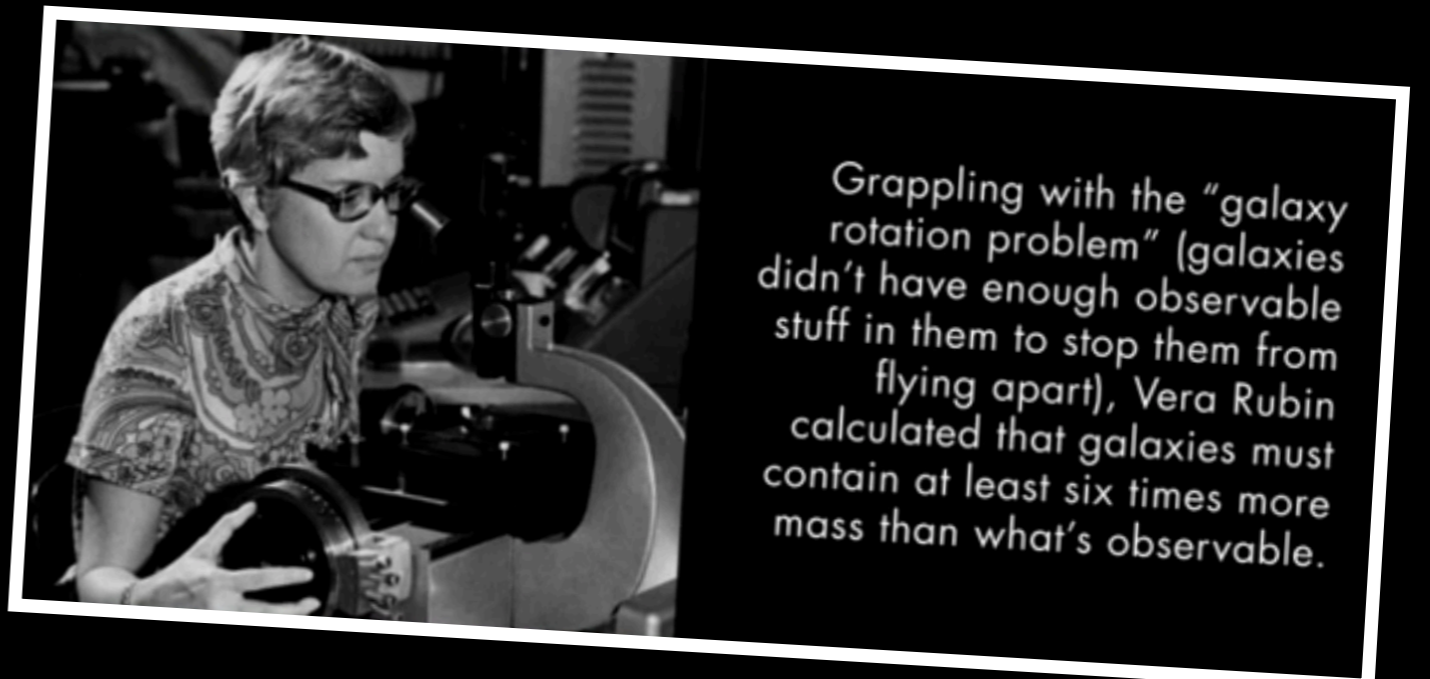
Part I: What have we learnt?

Part II: A new era in the quest for DM

Dark Matter “Mythology”



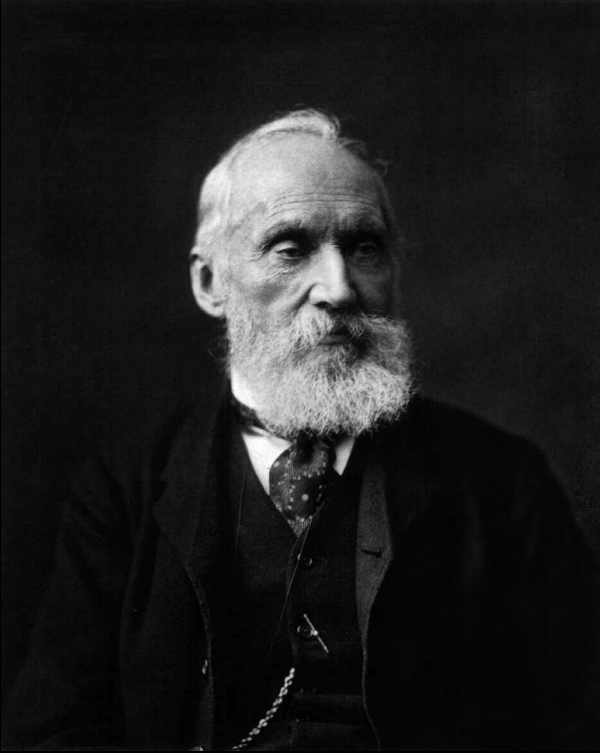
In 1933, Swiss astronomer Fritz Zwicky applied a mathematical theorem to infer the existence of what he called Dunkle Materie, coining the term dark matter. Zwicky was a noted curmudgeon and self-described “lone wolf” who claimed to “have a good idea every two years.”



Grappling with the “galaxy rotation problem” (galaxies didn’t have enough observable stuff in them to stop them from flying apart), Vera Rubin calculated that galaxies must contain at least six times more mass than what’s observable.

Figures: Perimeter Institute

Dark matter: a problem with a long history..



Lord Kelvin (1904)

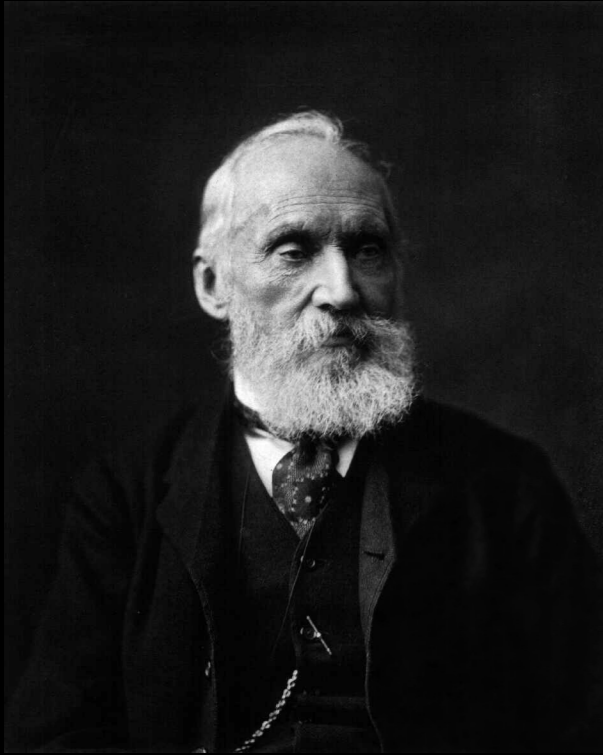
“Many of our stars, perhaps a great majority of them, may be dark bodies.”



Henri Poincaré (1906)

*“Since [the total number of stars] is comparable to that which the telescope gives, then there is no **dark matter**, or at least not so much as there is of shining matter.”*

Dark matter: a problem with a long history..



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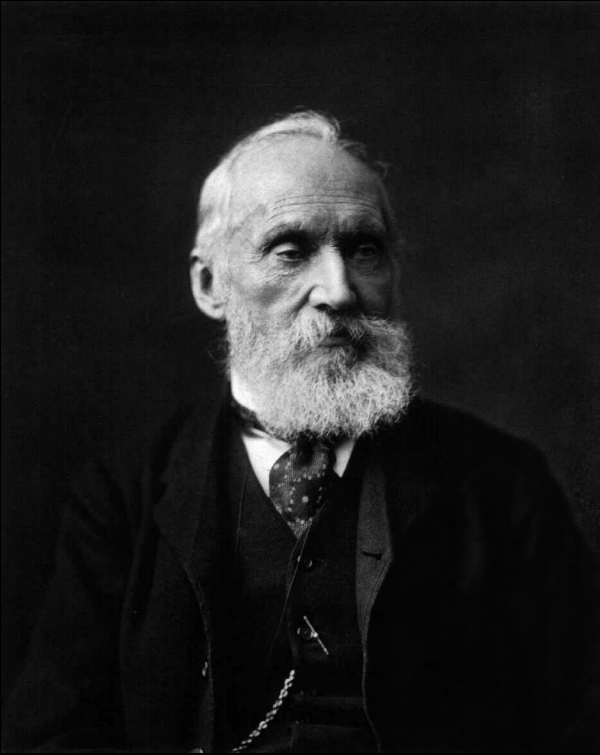
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Albert Einstein (1921)

Applies virial theorem to star cluster: “the non luminous masses contribute no higher order of magnitude to the total mass than the luminous masses”

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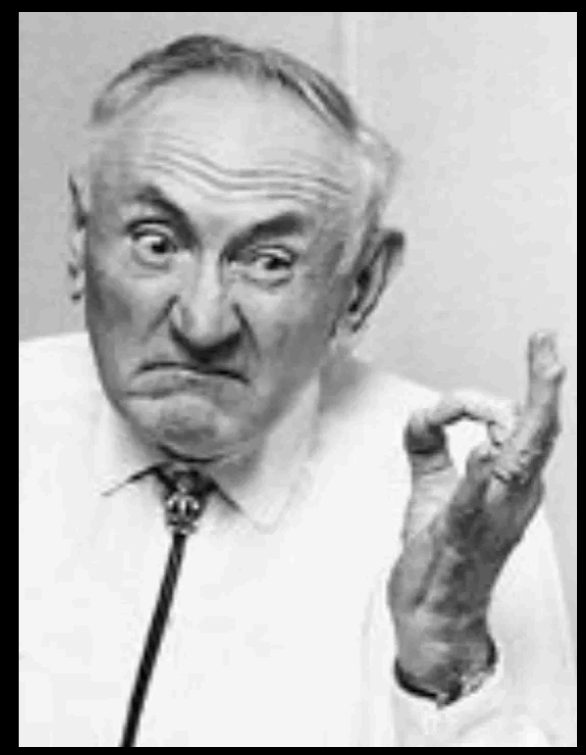
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Applies virial theorem to star cluster: “the non luminous masses contribute no higher order of magnitude to the total mass than the luminous masses”



Fritz Zwicky (1933)

“According to present estimates the average density of dark matter in our galaxy and throughout the rest of the universe are in the ratio 10^5 ”

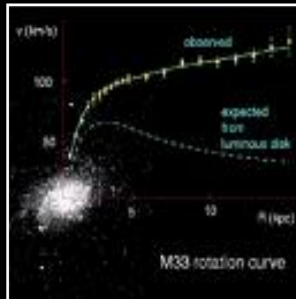
“A history of Dark Matter” GB & Hooper - RMP 1605.04909

“How dark matter came to matter” de Swart, GB, van Dongen - Nature Astronomy; 1703.00013

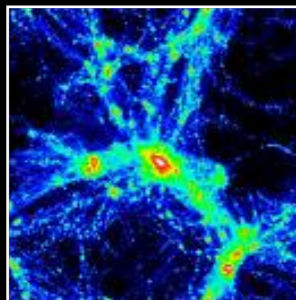


What is the Universe made of?

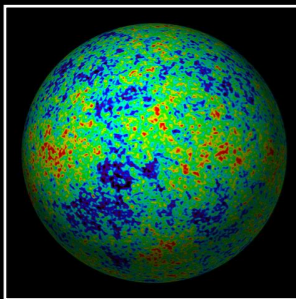
OBSERVATIONS



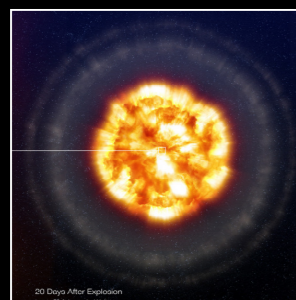
- Rotation Curves



- Clusters of galaxies

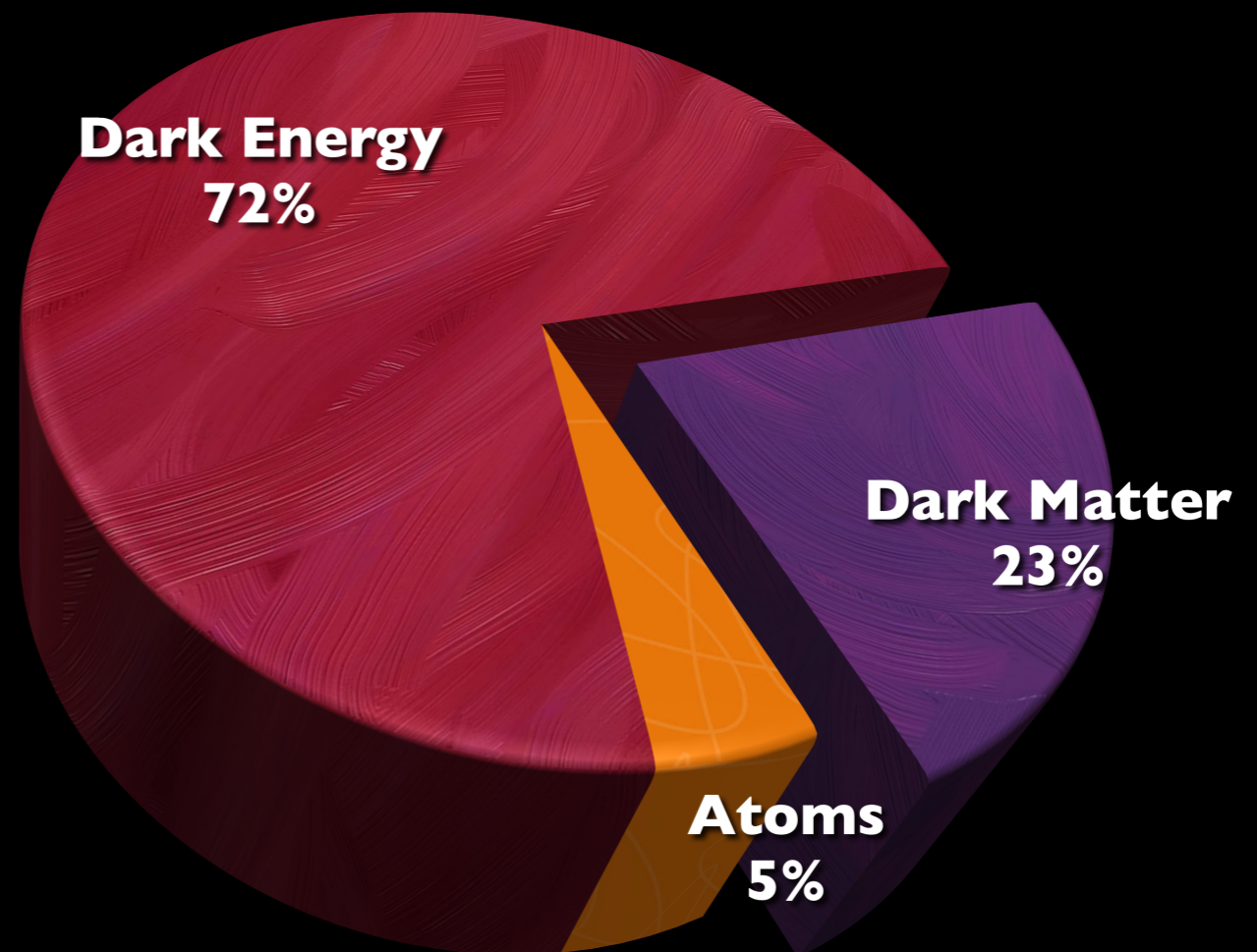


- CMB



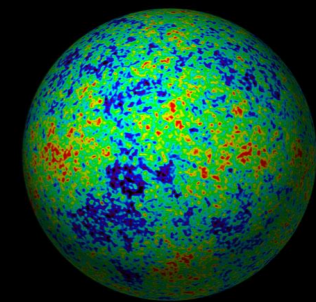
- Type Ia Supernovae

...

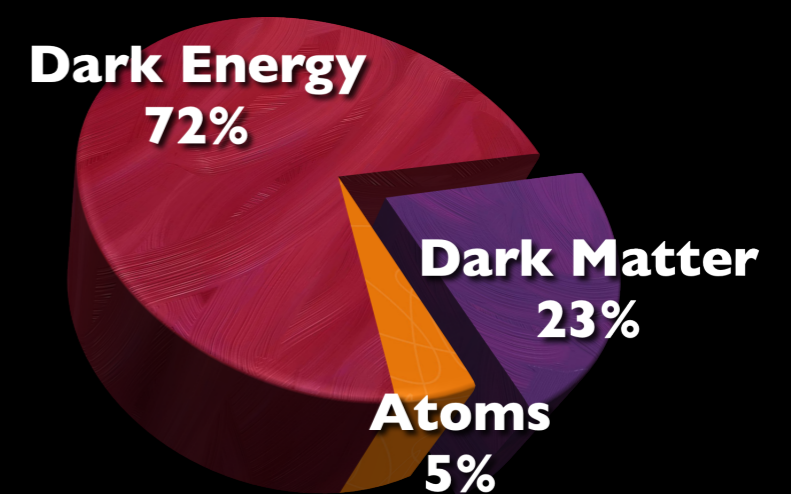
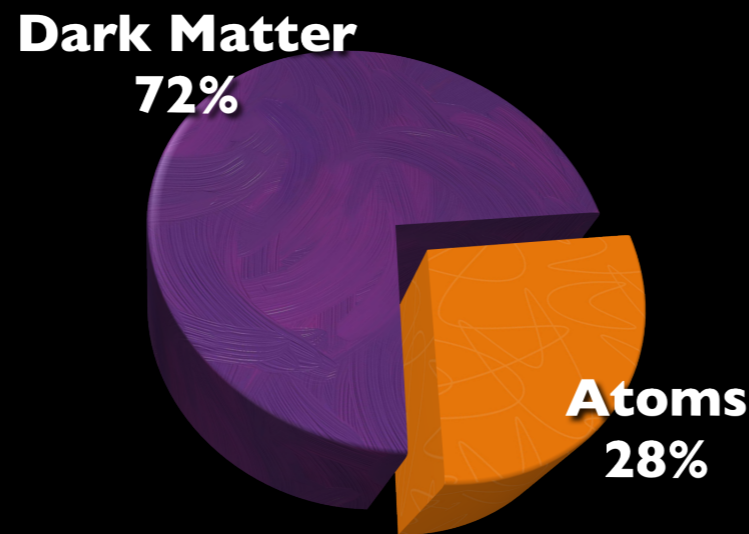


[statement valid now, and on very large scales]

What is the Universe made of?

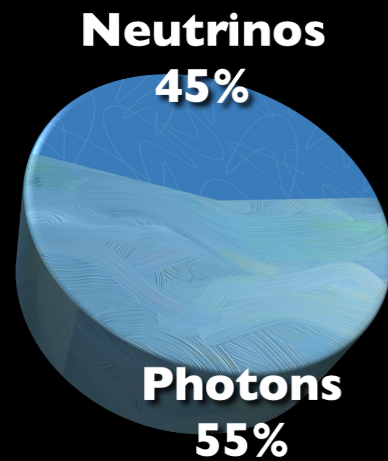


Posti & Helmi, A&A 621,A56 (2019)

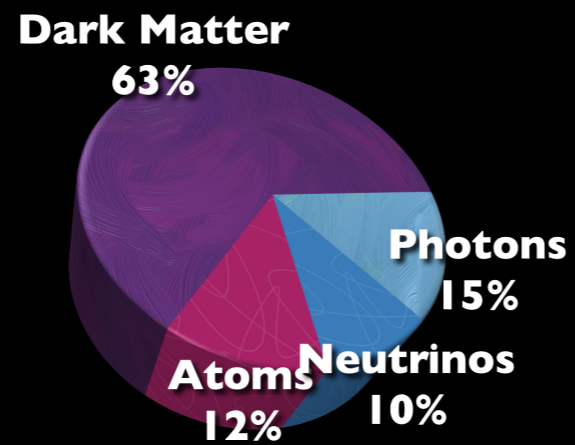


What was the Universe made of?

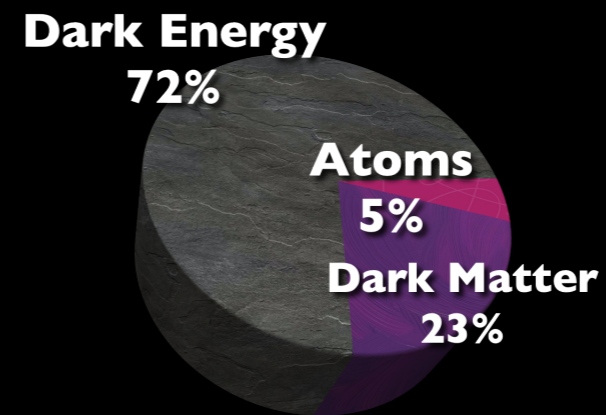
At BBN



At recombination



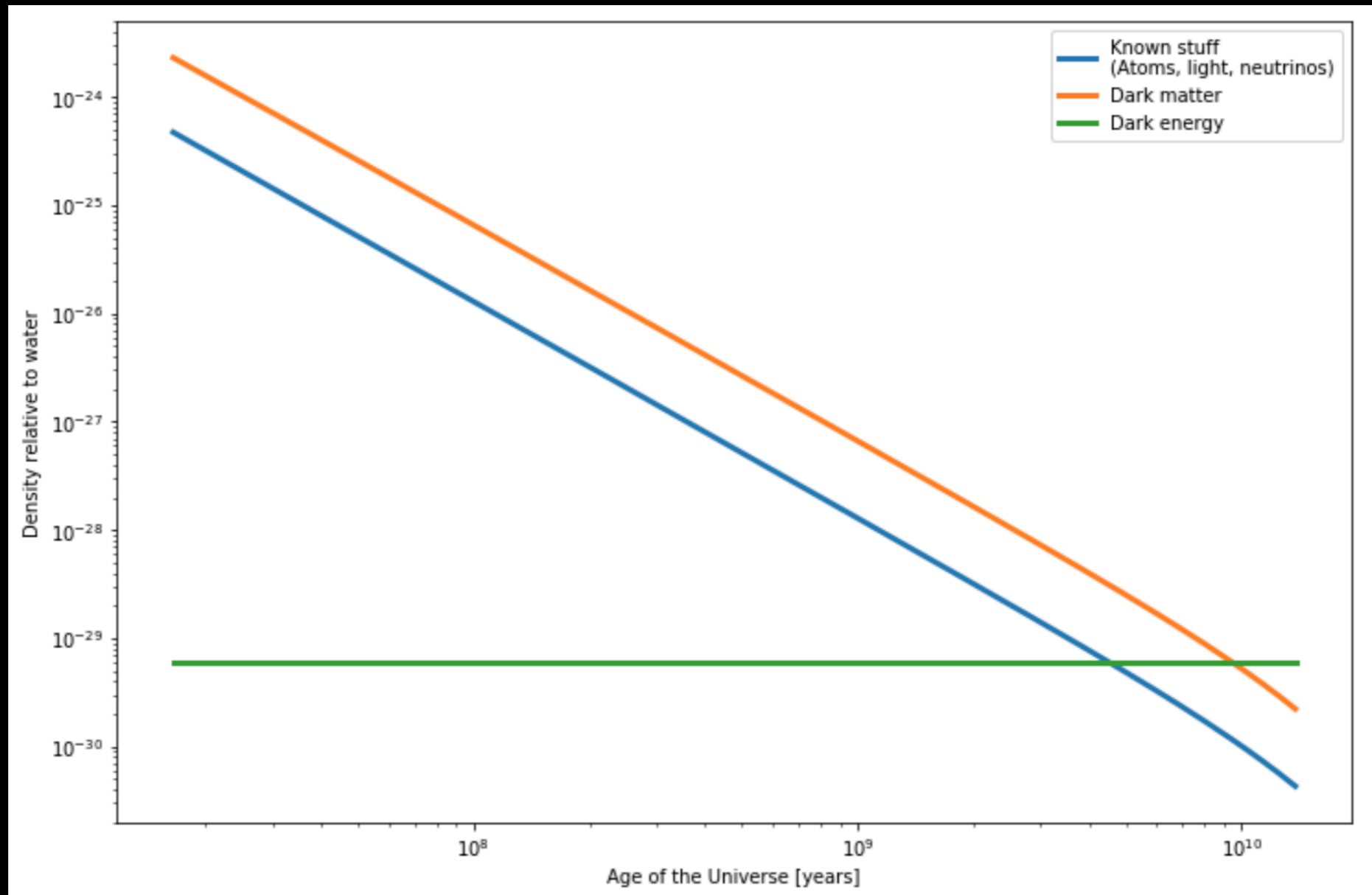
Today



...eventually



Evolution of matter/energy density



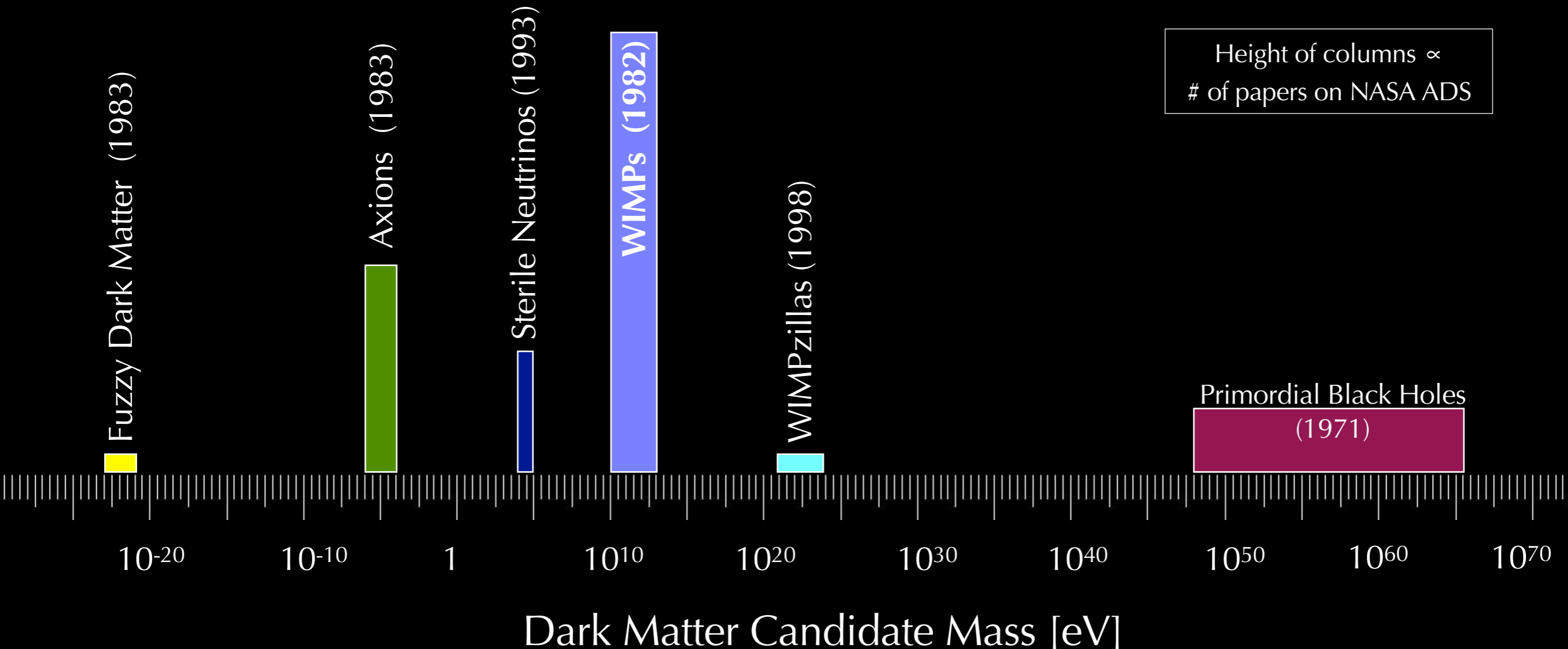
Created with #astropy <https://astropy.org>, astropy.cosmology package <https://docs.astropy.org/en/stable/cosmology/>

Candidates

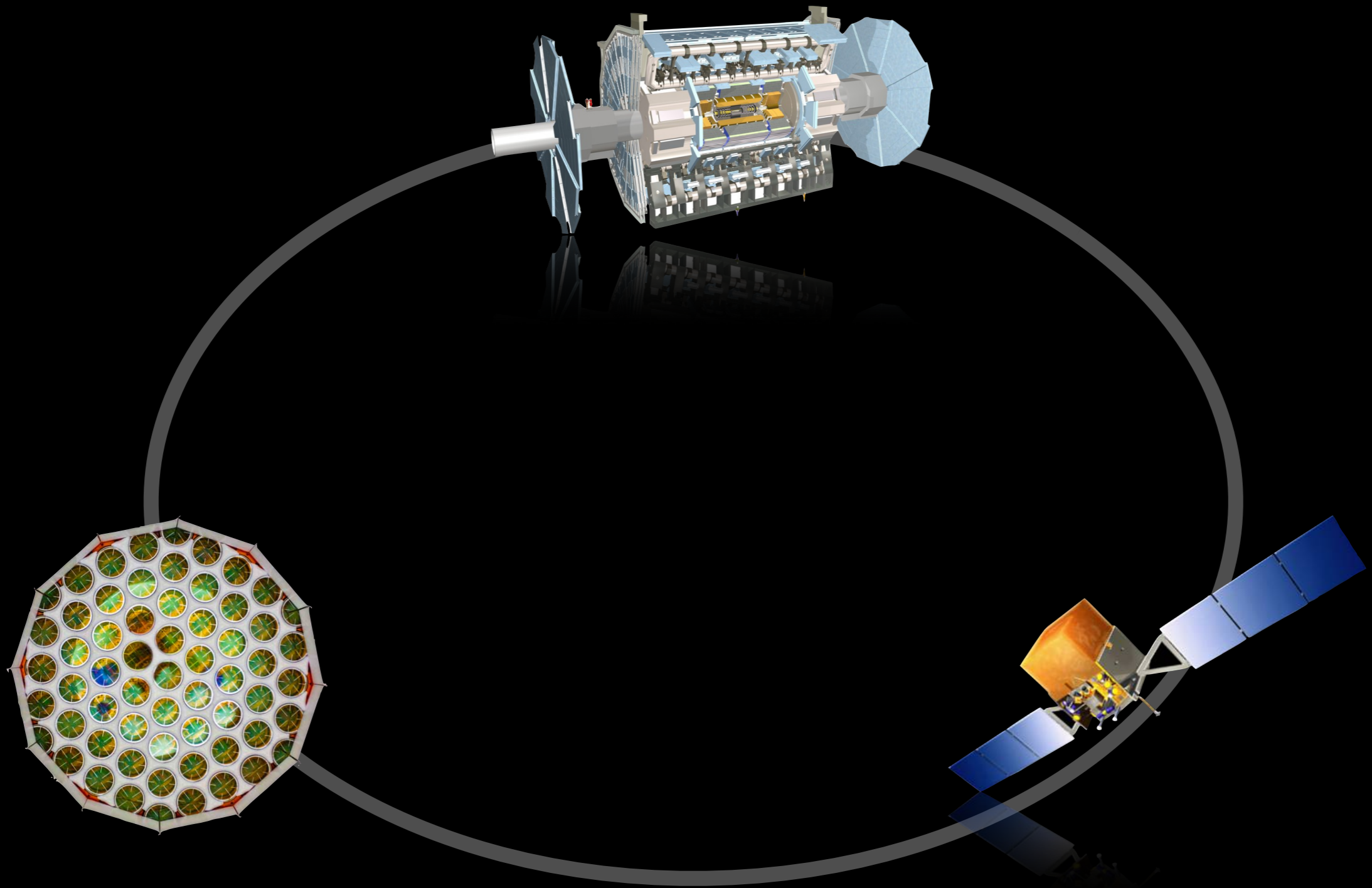


Candidates

- No shortage of ideas..
- Tens of dark matter models, each with its own phenomenology
- Models span 90 orders of magnitude in DM candidate mass!



WIMPs searches



WIMPs searches

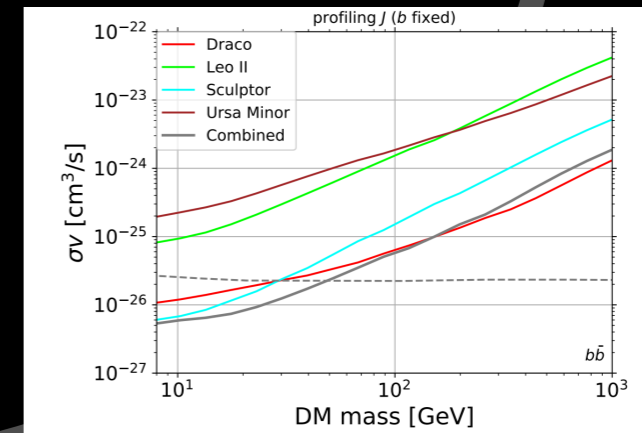
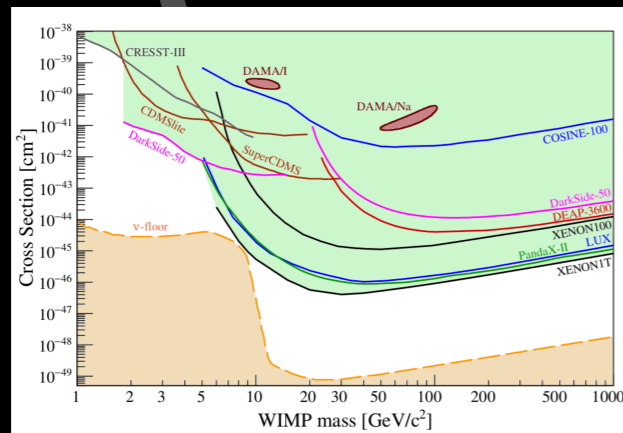
ATLAS SUSY searches

ATLAS SUSY Searches - 95% CL Lower Limits

Model	χ^2_{min}	$\chi^2_{min} + 1$	$\chi^2_{min} + 3.84$	Mass limit	Reference
CMSSM	1.0	1.0	1.0	1000	ATLAS Collaboration
gMSUGRA	1.0	1.0	1.0	1000	ATLAS Collaboration
UMSUGRA	1.0	1.0	1.0	1000	ATLAS Collaboration
CMSSM	1.0	1.0	1.0	1000	ATLAS Collaboration
gMSUGRA	1.0	1.0	1.0	1000	ATLAS Collaboration
UMSUGRA	1.0	1.0	1.0	1000	ATLAS Collaboration
CMSSM	1.0	1.0	1.0	1000	ATLAS Collaboration
gMSUGRA	1.0	1.0	1.0	1000	ATLAS Collaboration
UMSUGRA	1.0	1.0	1.0	1000	ATLAS Collaboration

ATLAS Preliminary
2017-12-13 13:30

No WIMPs
found yet, despite many efforts!



Are WIMPs ruled out?

NO

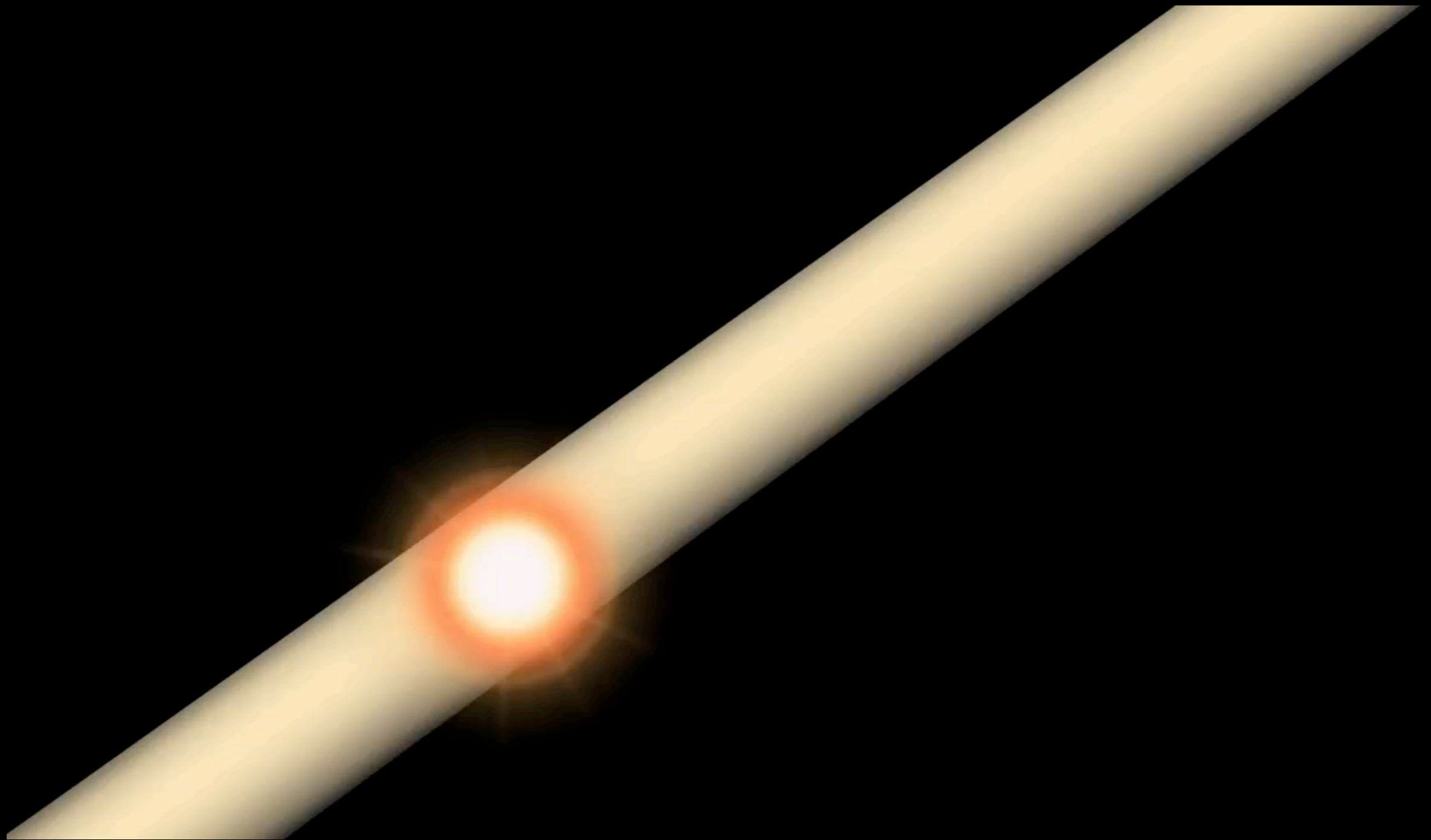
absence of evidence \neq evidence of absence

A new era in the search for DM

GB, Tait, *Nature* (2018) 1810.01668

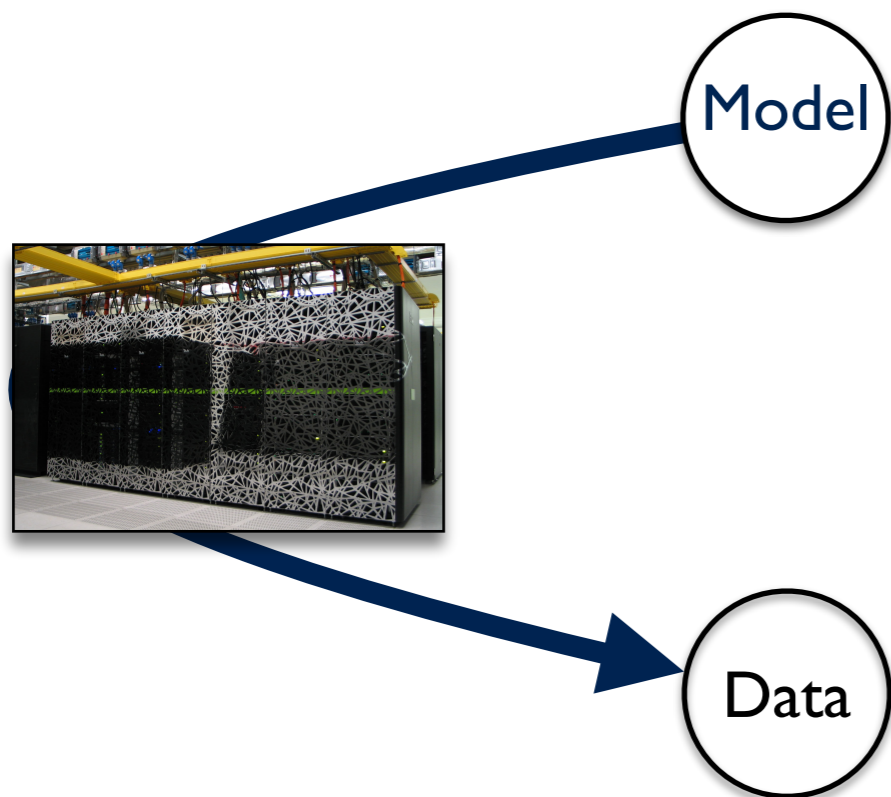
- I. Broaden/improve/diversify searches
- II. Exploit astro/cosmo observations
- III. Exploit Gravitational Waves

Dark matter searches at the LHC



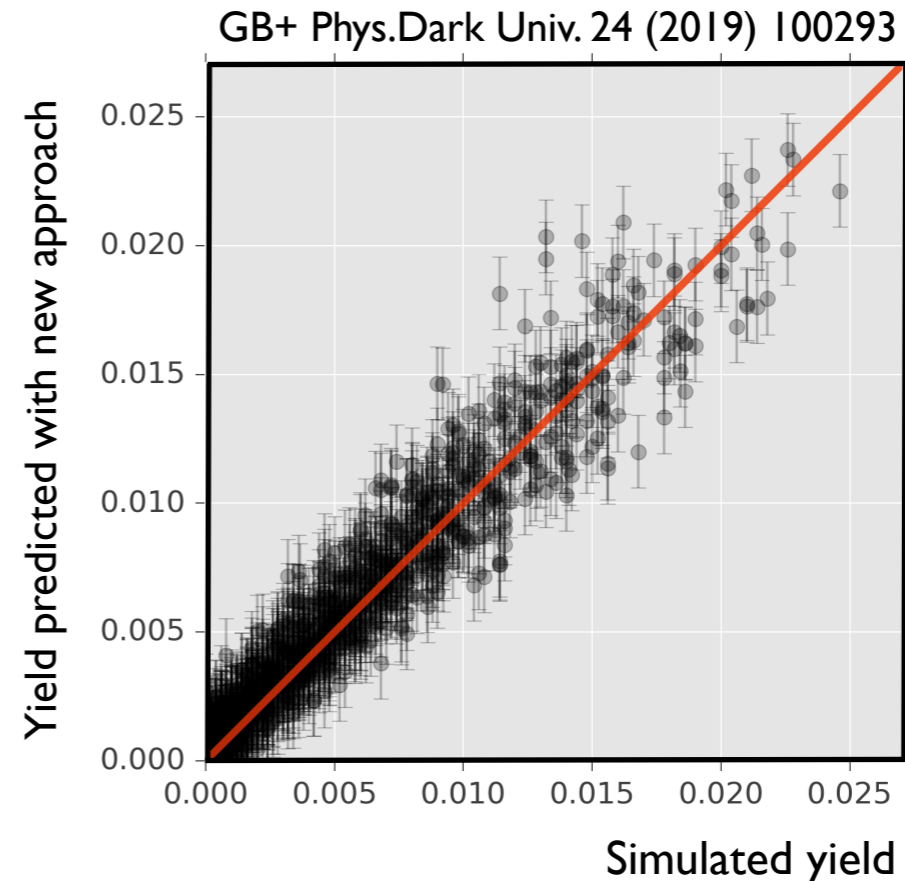
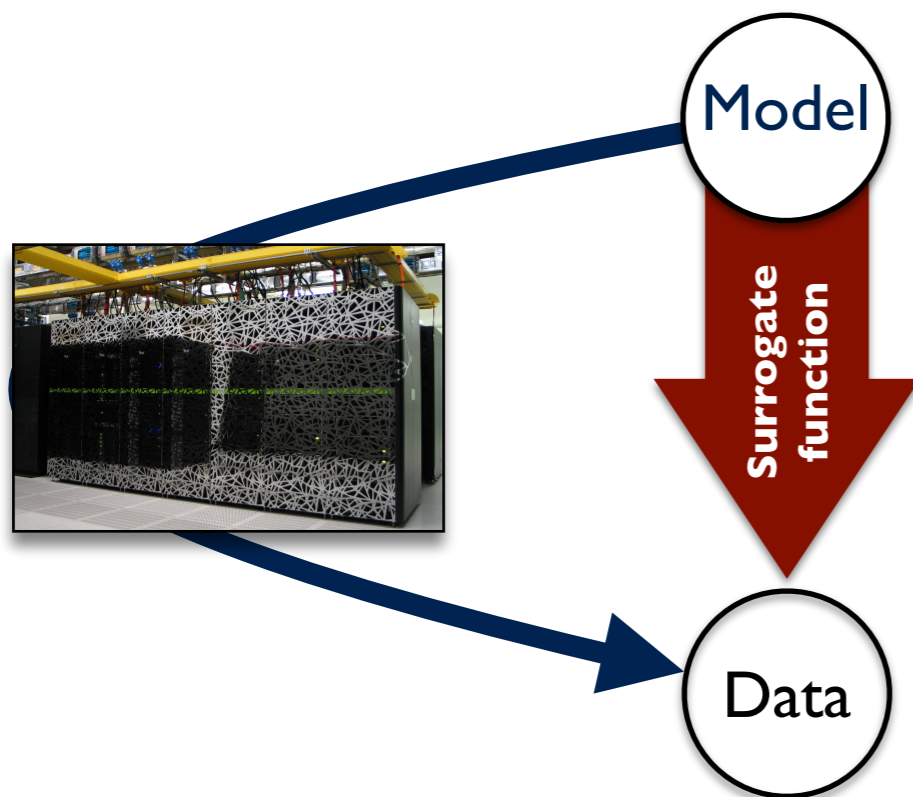
Improving existing strategies

Speeding up statistical inference with Machine Learning tools



Improving existing strategies

Speeding up statistical inference with Machine Learning tools

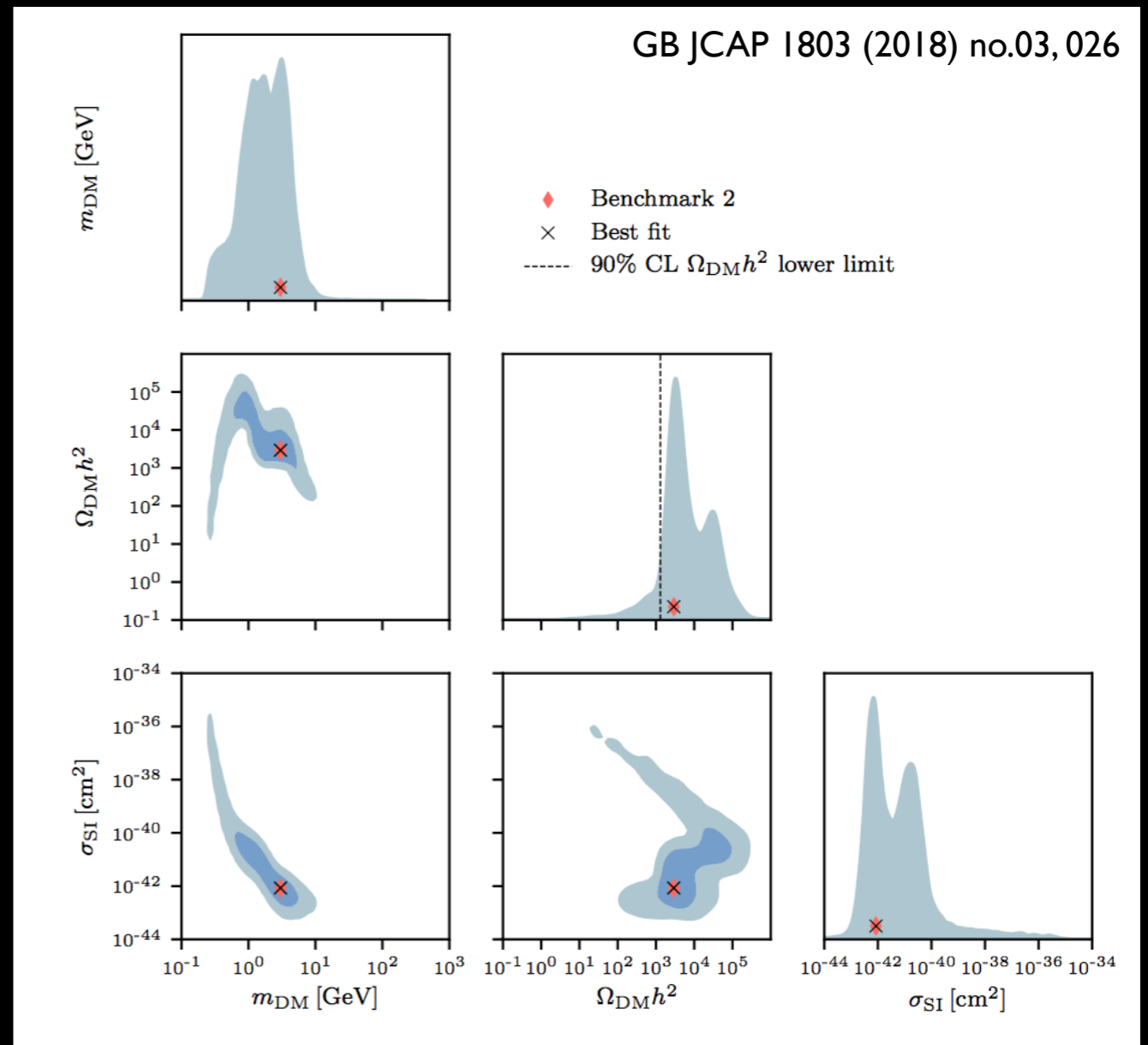


- Exploring parameter spaces of theoretical models computationally expensive
- Machine learning methods (*distributed gaussian processes, deep neural networks*) bring computation time from *~CPU centuries* to *~CPU weeks*!
- Can be run by a PhD student in 1 day on a desktop computer!

Improving existing strategies

E.g. New Machine Learning tools applied to LHC searches:

- i) Fast exploration of phenomenology in high-dimensional parameter spaces
- ii) Perform fast inference if new particles discovered, that allows us to recover theory parameters compatible with data



The *Dark Machines* initiative



About Dark Machines

Dark Machines is a research collective of physicists and data scientists. We are curious about the universe and want to answer cutting edge questions about Dark Matter with the most advanced techniques that data science provides us with.

3rd DarkMachines workshop: Advanced Workshop on Accelerating the Search for Dark Matter with Machine Learning

27 April 2020 to 1 May 2020

CERN

Europe/Zurich timezone

Postponed

Website: darkmachines.org ; Twitter: [dark_machines](https://twitter.com/dark_machines)

A new era in the search for DM

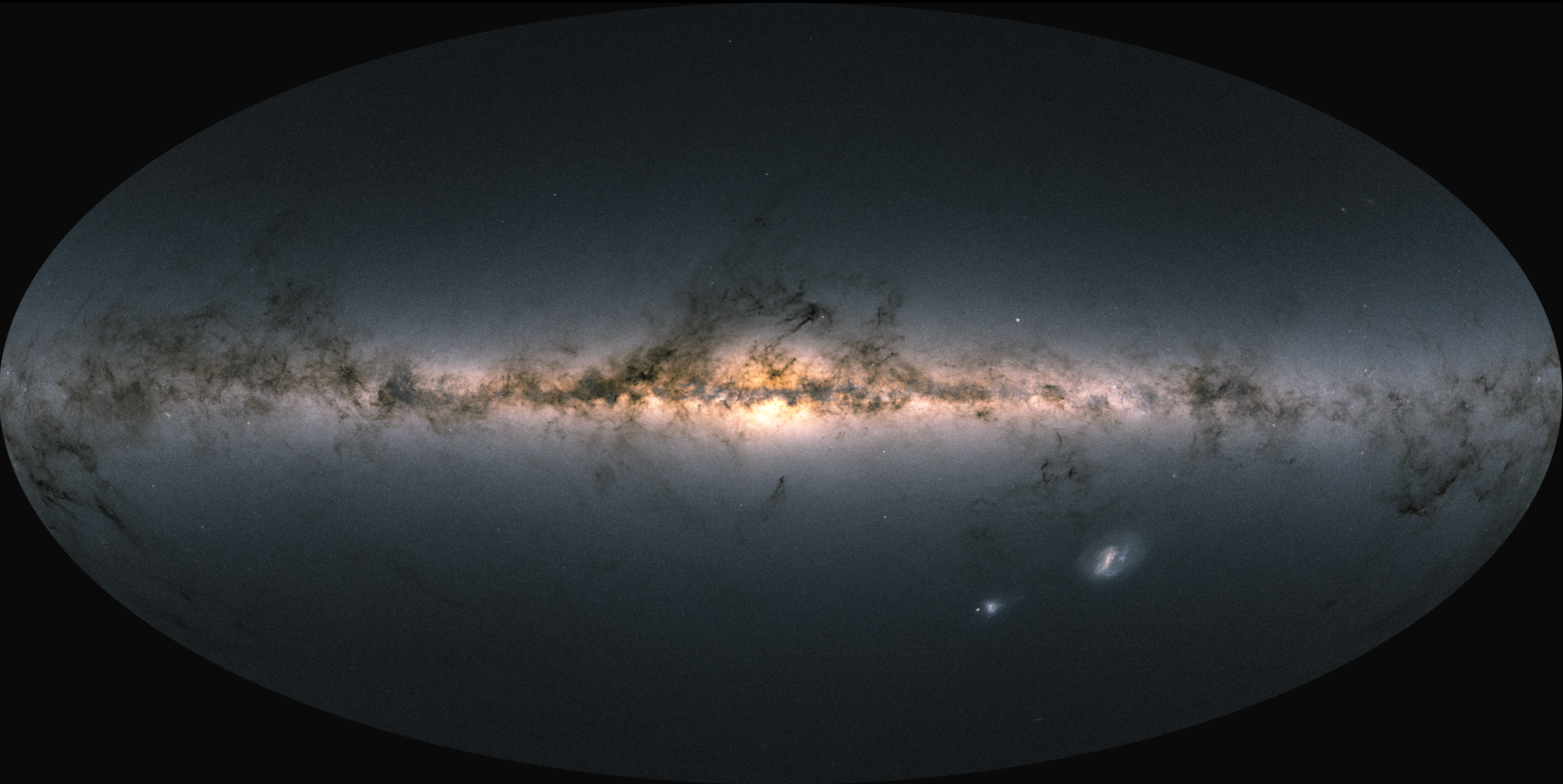
GB, Tait, *Nature* (2018) 1810.01668

I. Broaden/improve/diversify searches

II. Exploit astro/cosmo observations

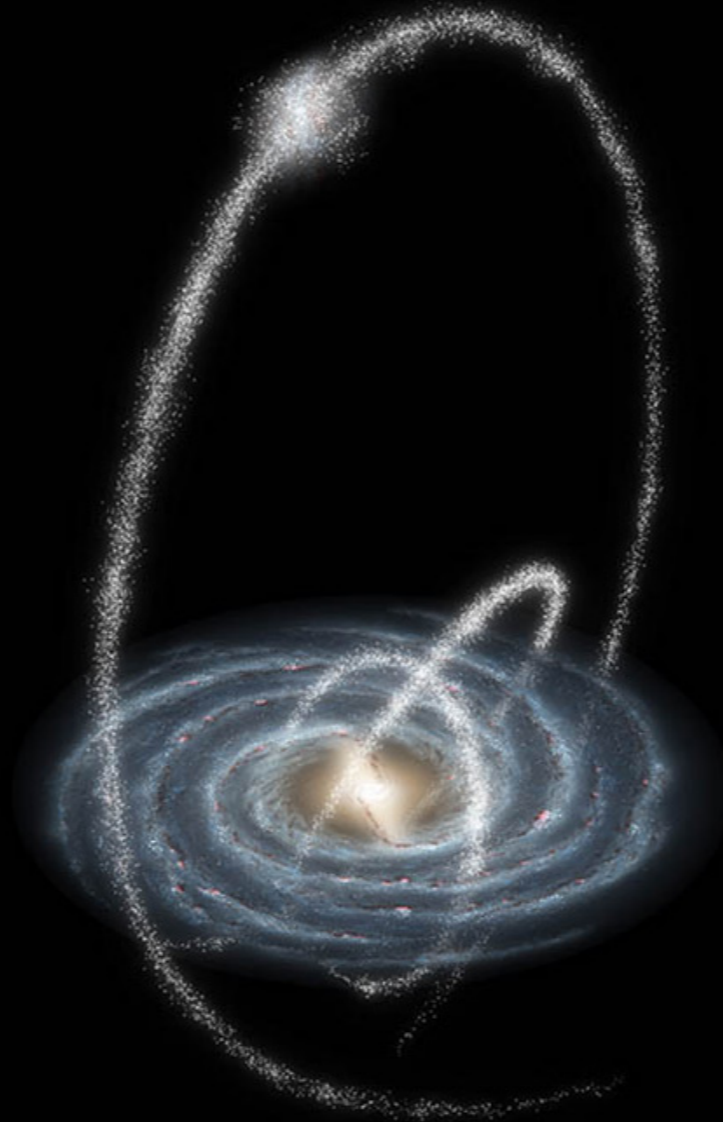
III. Exploit Gravitational Waves

GAIA'S SKY



Total brightness and colour of stars observed by ESA's Gaia satellite and released as part of Gaia's Early Data Release 3

Stellar streams



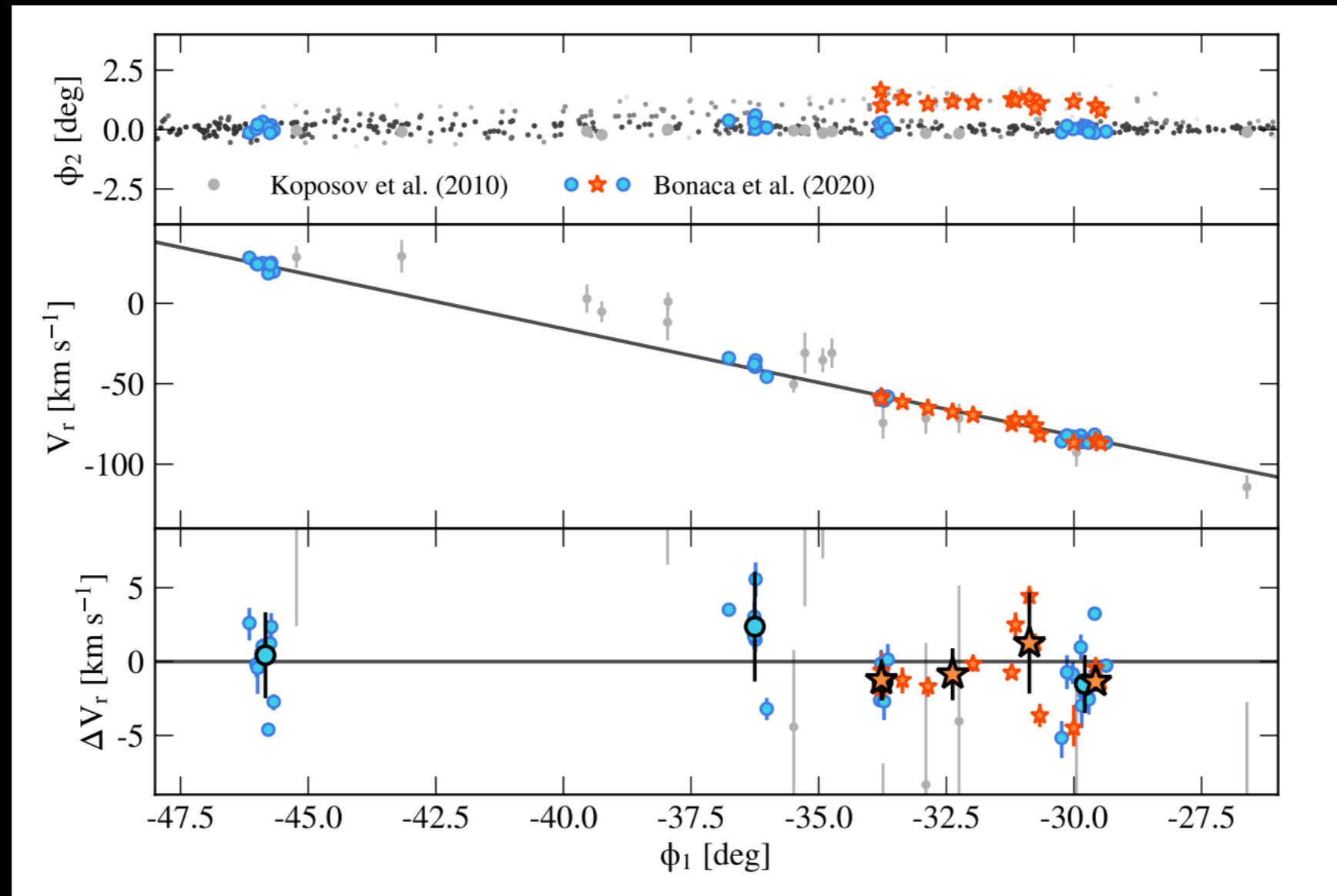
Searching for dark matter substructures in the MW



Gaia GDI stream data!

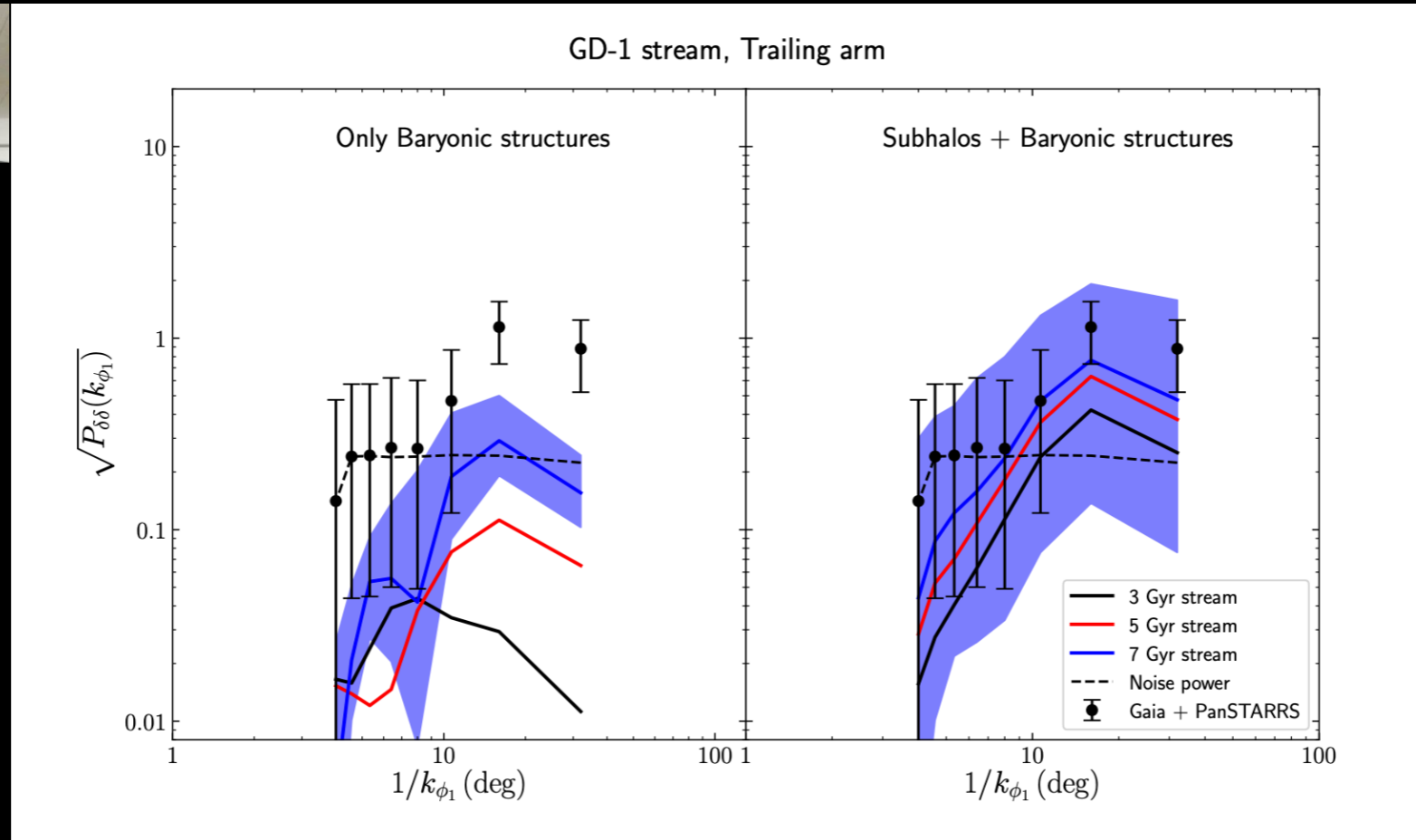
New map of stars in GDI stream (longest cold stream in the MW) with *Gaia* second data release combined with *Pan-STARRS*.

Stream appears to be perturbed, with several ‘gaps’ and a ‘spur’



Bonaca et al. 2001.07215

Statistical analysis of perturbations: Strong hints of dark substructures!



Banik, Bovy, GB, Erkal, de Boer, MNRAS 502, 2364 (2021)

- Gaia GD1 stream data exhibit substantial ‘structure’
- Density fluctuations cannot be explained by “baryonic” structures (GC, GMC, spiral arms etc)
- **Density fluctuations are consistent with CDM predictions (not a fit!)**

The future of dark matter searches

- I. Broaden/improve/diversify searches
- II. Exploit astro/cosmo observations
- III. Exploit Gravitational Waves

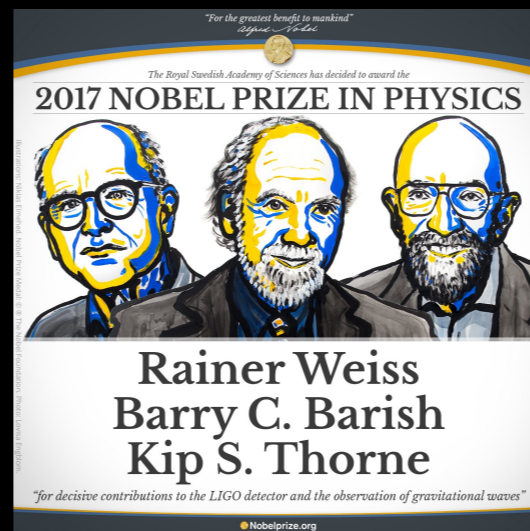
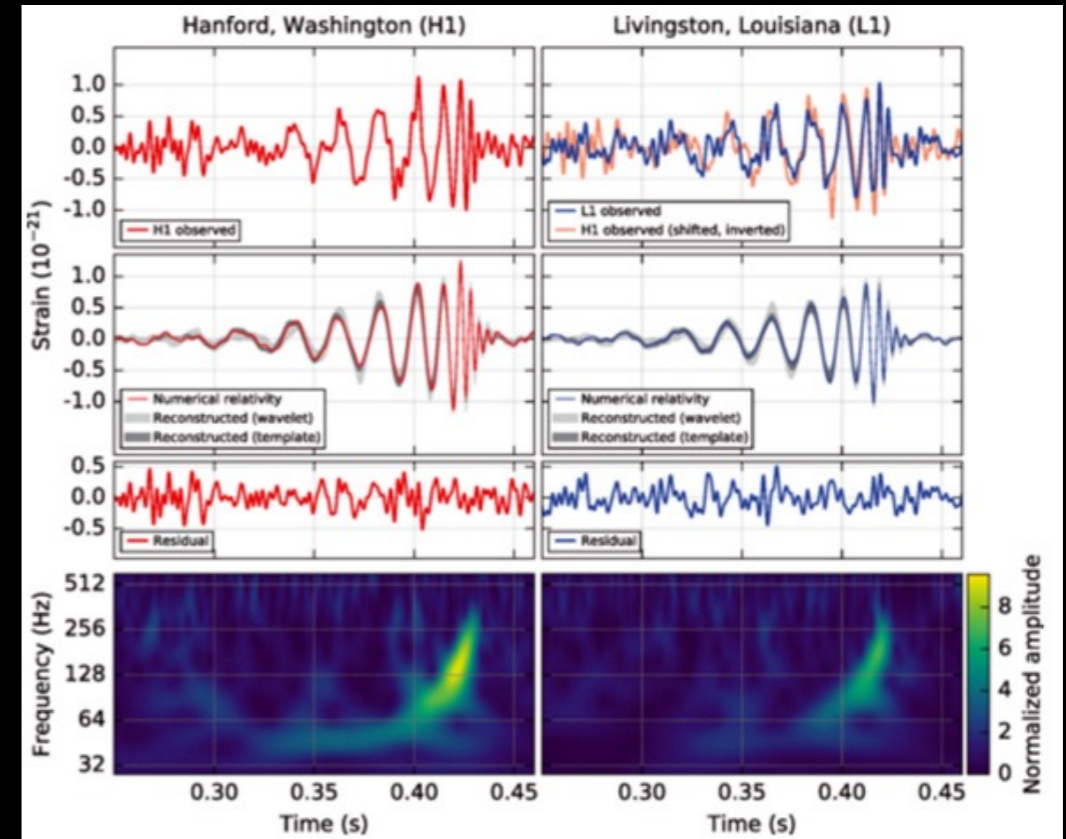
Gravitational Waves

“The discovery that shook the world”

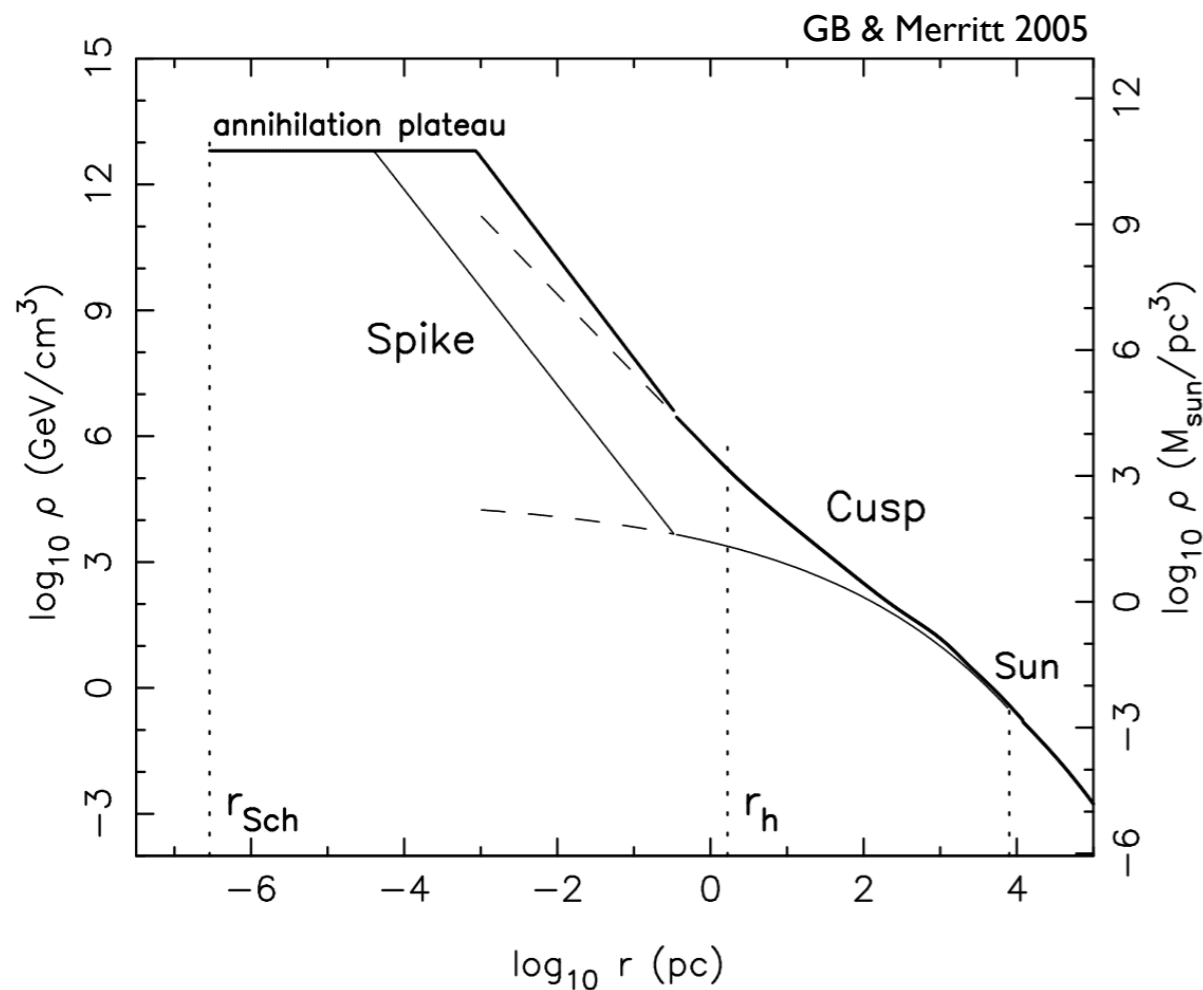
LIGO & Virgo coll, PRL 116, 061102



Primary black hole mass $36^{+5}_{-4} M_{\odot}$
Secondary black hole mass $29^{+4}_{-4} M_{\odot}$



Dark Matter ‘dress’ around BHs

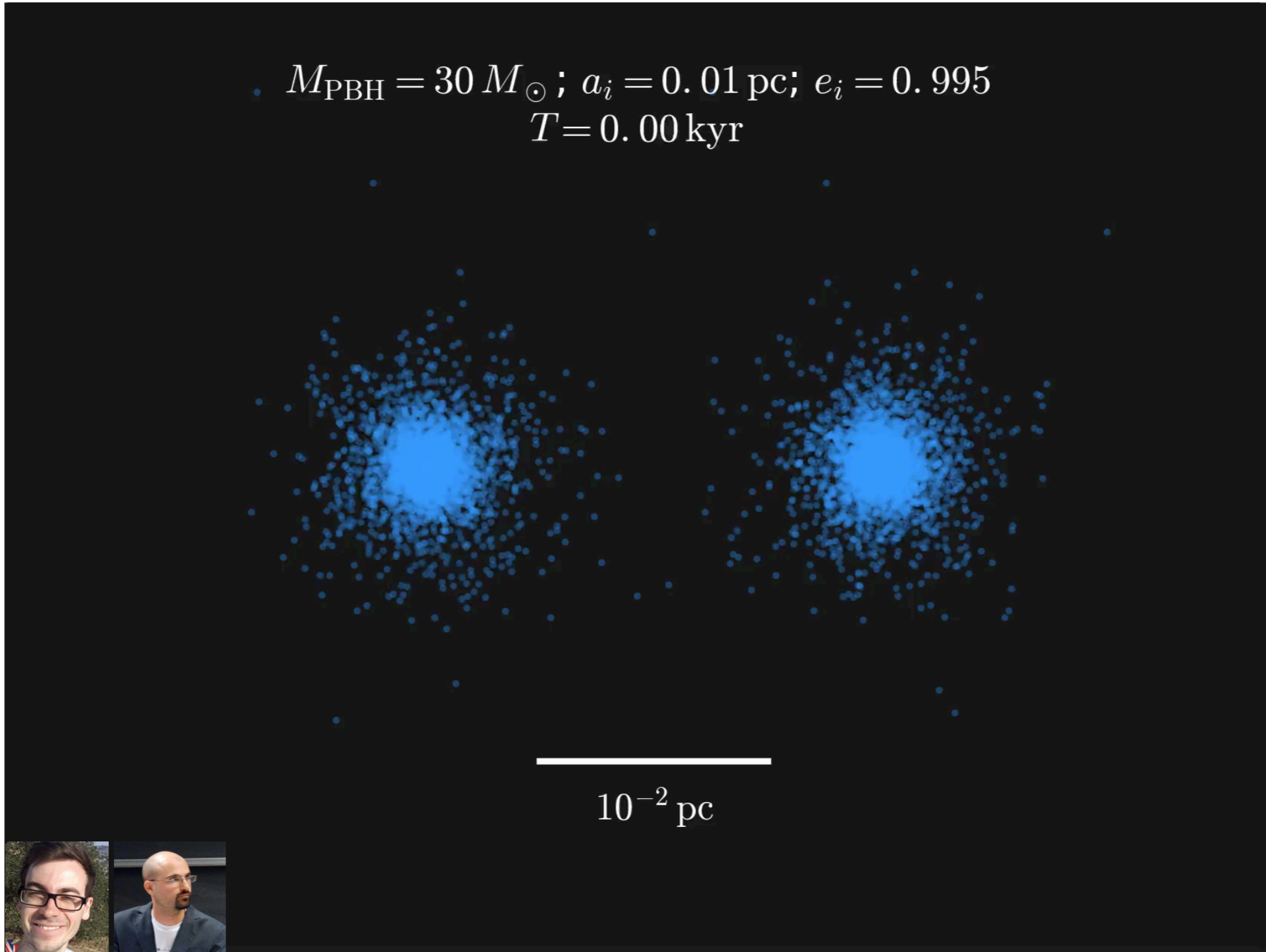


- **Adiabatic ‘spikes’ around SMBHs**
(Gondolo & Silk 2000)
- **‘Mini-spikes’ around IMBHs**
(GB, Zentner, Silk 2005)
- **Overdensities around primordial BHs**
(e.g. Adamek et al. 2019)
- **Ultralight boson ‘clouds’**
(e.g. Brito, Cardoso & Pani 2015)

Open questions: astrophysical uncertainties, dependence on DM properties (self-interactions, annihilations)

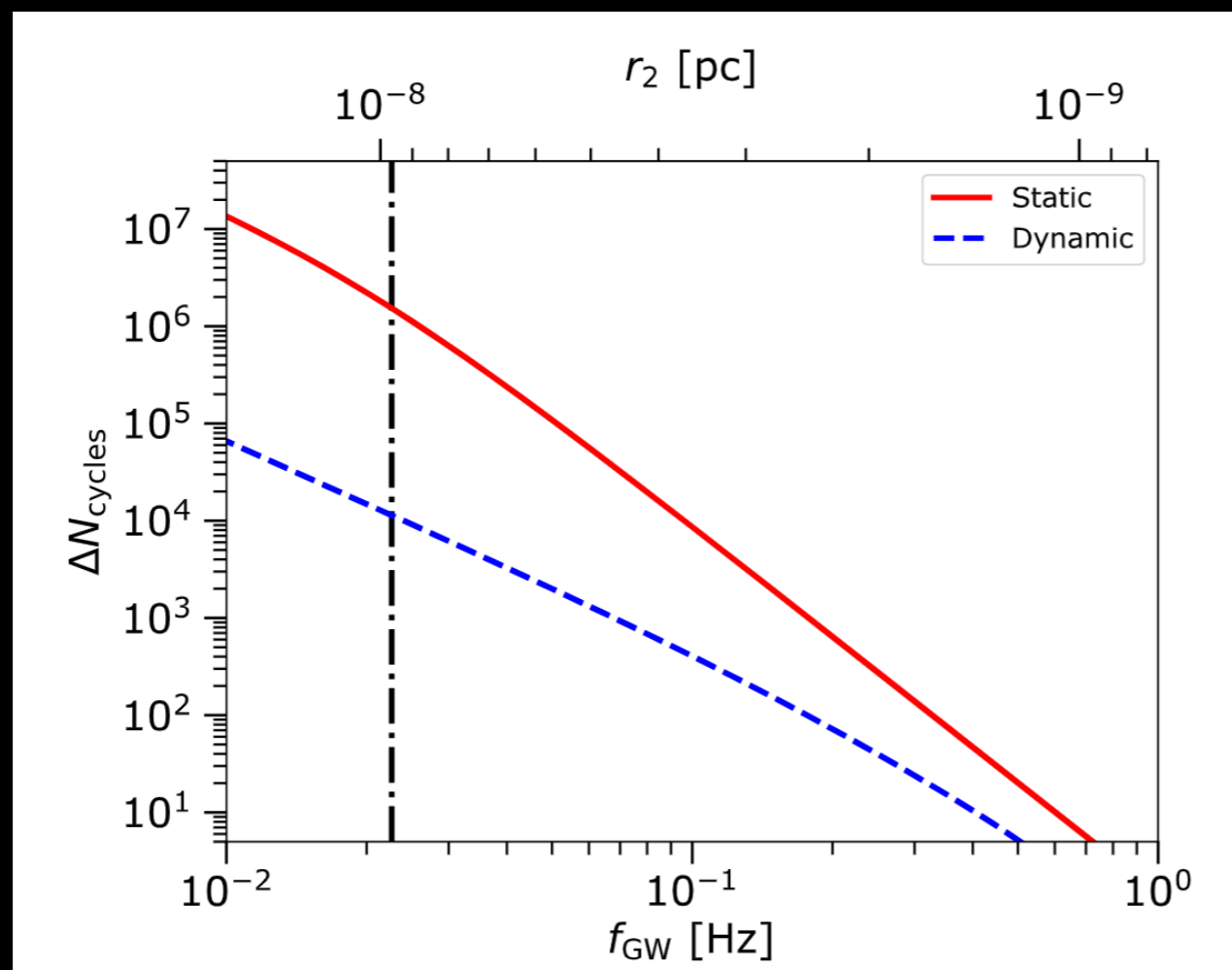
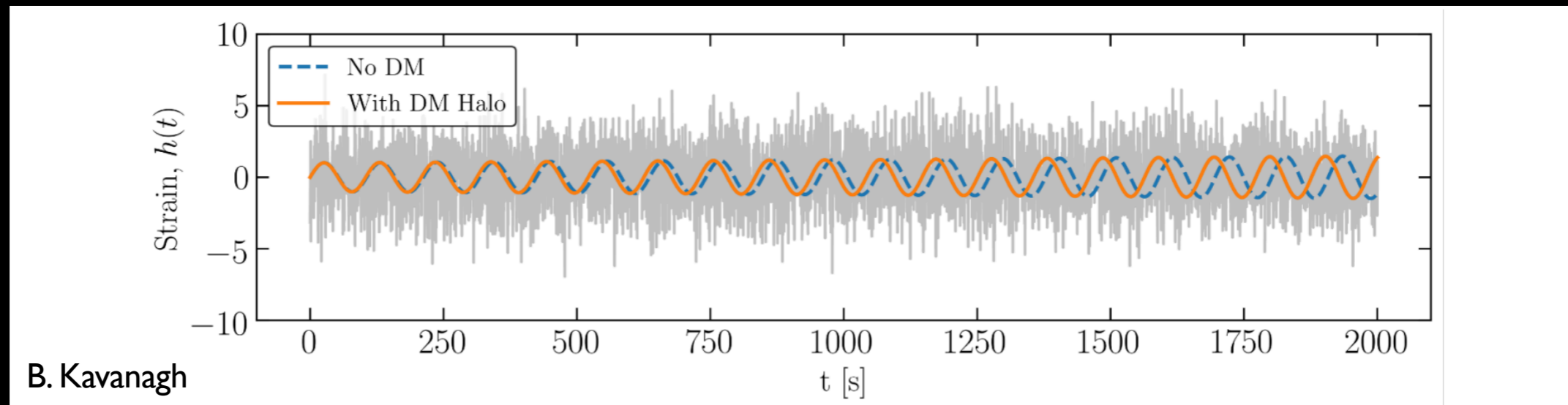
'Dressed' BH-BH merger

$$M_{\text{PBH}} = 30 M_{\odot}; a_i = 0.01 \text{ pc}; e_i = 0.995$$
$$T = 0.00 \text{ kyr}$$



Kavanagh, Gaggero & GB, arXiv:1805.09034

Gravitational Waveform dephasing



- Dark matter modifies binary dynamics via dynamical friction (Eda+ 2013, 2014)
- This induces a dephasing of the waveform, potentially detectable e.g. with LISA
- Dephasing is smaller than previously thought (i.e. wrt to case with fixed dark matter profile) but still potentially detectable

Conclusions

- This is a time of profound transformation for dark matter studies, in view of the absence of evidence (though NOT evidence of absence) of popular candidates
- LHC, ID and DD experiments may still reserve surprises!
- At the same time, it is urgent to:
 - Diversify dark matter searches
 - Exploit astronomical observations
 - Exploit gravitational waves
- The field is completely open: extraordinary opportunity for new generation to come up with new ideas and discoveries